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# Settling velocity of small inertial particles in homogeneous isotropic turbulence from high-resolution DNS



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# ABSTRACT

The gravitational settling velocity of small heavy particles in a three-dimensional turbulent flow remains a controversial topic. In a homogeneous turbulence of zero mean velocity, both enhanced settling velocity and reduced settling velocity have been reported relative to the still-fluid terminal velocity. Dominant mechanisms for enhanced settling include the preferential sweeping and particle-particle hydrodynamic interactions. The reduced settling could result from loitering (falling particles spend more time in the regions with upward flow), vortex trapping, and drag nonlinearity. Here high-resolution direct numerical simulations (DNS) are used to investigate the settling velocity of non-interacting small heavy particles, for an extended range of flow Taylor microscale Reynolds numbers (up to  $R_{\lambda} = 500$ ) with varying particle terminal velocity (relative to the Kolmogorov velocity) and particle inertia, by changing the particle-tofluid density ratio and energy dissipation rate. For the parameter regimes considered here, the preferential sweeping has a dominant effect leading to an increase of the average settling velocity relative to the terminal velocity; and this increase is mainly governed by particle Froude number (the ratio between the particle inertial response time and the residence time of the particle in a Kolmogorov eddy) and its magnitude depends linearly on the square root of the energy dissipation rate. The reduction of settling due to loitering rarely occurs in a homogeneous turbulence without organized large-scale vortical structures, but is found to emerge only if the particle horizontal motions are blocked (thus removing the preferential sweeping effect), as shown in Good et al. (2014). The DNS results were used to develop a parameterization that relates the settling velocity to the particle inertia (St), Froude number, and  $R_{\lambda}$ . Finally, sensitivities of the DNS results to the large-scale forcing method and to the drag nonlinearity are also briefly discussed.

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# 1. Introduction

The settling velocity of small heavy particles is relevant to many applications such as pipeline pneumatic transport (Fokeer et al., 2004; Lain and Sommerfeld, 2013), coal combustion (Zhou et al., 2004; Smoot, 2013), transport of biogenic substances in the oceans (Noh et al., 2006), sediment transport in water bodies (Papanicolaou et al., 2008; Keshtpoor et al., 2015), modeling of dust storm (Tsidulko et al., 2002), and warm rain formation (Grabowski and Wang, 2013; Ghosh and Jonas, 2001). More generally, the dynamics of small heavy particles in a turbulent flow affects the

http://dx.doi.org/10.1016/j.ijmultiphaseflow.2016.04.005 0301-9322/© 2016 Elsevier Ltd. All rights reserved. spatial distribution of particles (Squires and Eaton, 1991; Wang and Maxey, 1993), particle deposition rate (Kallio and Reeks, 1989; Li et al., 2001), and turbulent collision rate of inertial particles (Zhou et al., 2001; Sundaram and Collins, 1997).

Small heavy particles have a diameter less than the Kolmogorov scale of the carrier fluid turbulence and a density much larger than that of the fluid. These particles could have a significant inertia, as quantified by a finite Stokes number *St* (the ratio of their inertial response time  $\tau_p$  to the flow Kolmogorov time  $\tau_K$ ), and a significant terminal velocity, in terms of the ratio  $S_V = V_T/v_K$  with  $V_T$  being the particle still-fluid terminal velocity and  $v_K$  the flow Kolmogorov velocity. In this work, we concern mainly the average settling velocity of such particles in a homogeneous turbulent flow, under the assumption that the particle mass loading is very low so that the effect of particles on the fluid turbulence is weak (i.e.,

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 Table 1

 Relations between dimensional and nondimensional parameters.

	St	S <sub>V</sub>	Fr	$R_{\lambda}$
$a \\ g \\ \varepsilon \\ v \\ \sigma = \frac{\rho_p}{\rho_f} $	$St \propto a^{2}$ no effect $St \propto \varepsilon^{0.5}$ $St \propto \nu^{-1.5}$ $St \propto \sigma$ no effect	$S_V \propto a^2$ $S_V \propto g$ $S_V \propto \varepsilon^{-0.25}$ $S_V \propto v^{-1.25}$ $S_V \propto \sigma$ no effect	$Fr \propto a^{6}$ $Fr \propto g^{2}$ no effect $Fr \propto v^{-4}$ $Fr \propto \sigma^{3}$ no effect	no effect no effect $R_{\lambda} \propto \varepsilon^{-0.5}$ $R_{\lambda} \propto \nu^{-0.5}$ no effect $R_{\lambda} \propto N^{2/3}$

one-way coupling) and particle-particle local hydrodynamic interactions can be neglected. It is well known that the average settling velocity, relative to the mean flow, may differ from  $V_T$  for a variety of reasons, to be discussed below.

The motion of small heavy particles depends on two governing parameters which can be expressed in terms of other physical parameters as follows (Wang et al., 2006):

$$St = \frac{\tau_p}{\tau_K} = \frac{2}{9} \frac{\rho_p}{\rho_f} \frac{\varepsilon^{0.5}}{\nu^{1.5}} a^2, \quad S_V = \frac{V_T}{\nu_K} = \frac{2}{9} \frac{\rho_p}{\rho_f} \frac{g}{\varepsilon^{0.25} \nu^{1.25}} a^2, \tag{1}$$

which implies that, in real applications, there are five physical parameters affecting the dynamics: the particle radius *a*, the gravitational acceleration **g**, the energy dissipation rate  $\varepsilon$ , the fluid kinematic viscosity  $\nu$  and the particle-to-fluid density ratio  $\sigma = \rho_p / \rho_f$ .

An important derived parameter combining St and  $S_V$  is the particle Froude number, namely, the ratio of particle response time  $\tau_p$ to the residence time of the particle in a Kolmogorov eddy (Davila and Hunt, 2001)

$$Fr = \frac{\tau_p}{\Gamma_{vort}/\nu_p^2} = \frac{\tau_p V_T^2}{\eta \nu_K} = St S_V^2 = \frac{\tau_p^3 g^2}{\nu} = \left(\frac{2}{9}\right)^3 \left(\frac{\rho_p}{\rho_f}\right)^3 \frac{g^2 a^6}{\nu^4}, \quad (2)$$

which is independent of flow energy dissipation rate. Here,  $\eta$  is the Kolmogorov length scale. For water droplets in air, this parameter reaches one when  $a \approx 20 \ \mu$ m. Ayala et al. (2008) showed that the sedimenting particles obtain a maximum relative settling-velocity enhancement at  $a \approx 20 \ \mu$ m (see also Table 4 in Section 3.1). Another parameter affecting the level of settling velocity is the flow Reynolds number ( $R_{\lambda}$ ) which defines the range of flow length scales or the large-to-small scale ratio in the flow.  $R_{\lambda} = u'\lambda/\nu$  where u' is the rms fluctuating velocity and  $\lambda$  is the transverse Taylor microscale. The effects of key parameters on the dimensionless parameters are shown in Table 1.

Based on the simplified equation of motion we assume that the difference between average settling velocity  $\langle V_S \rangle$  and the terminal velocity  $V_T$  is an unknown function of the following parameters

$$\langle V_S \rangle - V_T = \mathbf{F}(\tau_p, V_T, \varepsilon, u', L_f, \ldots),$$
(3)

where,  $L_f$  is the integral lengthscale of the turbulent flow. Since  $L_f \sim \eta R_{\lambda}^{1.5}$  and  $R_{\lambda} = \sqrt{15} (u'/v_K)^2$  the integral lengthscale can be omitted from Eq. 3. The dimensional analysis leads to three equivalent representations of Eq. 3, namely

$$\frac{\langle V_S \rangle - V_T}{\nu_K} = \mathbf{F}_1(St, S_V, R_\lambda, ...), \tag{4}$$

$$\frac{\langle V_S \rangle - V_T}{V_T} = \mathbf{F}_2(St, S_V, R_\lambda, ...) = \frac{\mathbf{F}_1}{S_V}, \tag{5}$$

$$\frac{\langle V_{\rm S}\rangle - V_{\rm T}}{u'} = \mathbf{F}_3(St, S_V, R_\lambda, ...) = \frac{15^{0.25}}{R_\lambda} \mathbf{F}_1.$$
(6)

In each of these three equations, the change in the settling rate can be expressed in terms of the same function  $\mathbf{F}_1$ . This function, depends only on the dimensionless parameters such as *St*, *S*<sub>V</sub> (or *Fr*), and *R*<sub> $\lambda$ </sub>.

In a homogeneous turbulence of zero mean velocity, both enhanced settling velocity and reduced settling velocity have been reported. The dominant mechanism for enhanced settling velocity of small heavy particles in a turbulent flow is the preferential sweeping (Wang and Maxey, 1993). In their pioneering work, Maxey and Corrsin (1986) demonstrated that inertial particles can distribute very non-uniformly in a steady non-uniform flow, and consequently sedimenting particles may converge to a path located on the downward side of a vortex, leading to a higher settling velocity. Maxey (1987) extended the above study to a time-dependent non-uniform flow, and found both analytically and numerically that inertial particles accumulate in regions of low vorticity and high strain rate, a phenomenon now known as the preferential concentration. He found that the preferential concentration leads to a bias of particle trajectory and an enhanced settling rate. Wang and Maxey (1993) performed direct numerical simulations to show that the preferential concentration is strongest when St and  $S_V$  are of the order one in a realistic turbulent flow. They also demonstrated that the mean particle settling velocity is significantly enhanced under similar conditions, with the level of enhancement depending on the large-scale flow statistics.

Fung (1997) modeled the motion of small spherical particles in an infinite, two-dimensional unsteady flow. He showed that enhancement of the settling velocity occurs when  $V_T/\sigma_u \leq 0.7$ . Here  $\sigma_u$  is a characteristic velocity of the flow. The maximum increase occurs when  $V_T \approx 0.5\sigma_u$ . In the steady flow (flow with one length scale), however, the total suspension takes place when  $V_T$  is of order of the characteristic fluid velocity. Lillo et al. (2008) studied the settling velocity of heavy particles in two-dimensional turbulent and laminar flows. They showed that at large flow velocities particles are effectively guided to the down-flow regions and the actual settling velocity is larger than  $V_T$ . Afonso (2008) investigated the settling of inertial particles in 2D cellular flow. He found that for square, static cellular flows and at relatively small St the settling velocity, is larger than the terminal velocity. However, starting from a certain critical value of St, the falling velocity becomes lower than  $V_T$ . Although the author does not state explicitly what mechanism reduces the settling velocity, we assume that the reduction may be due to vortex trapping.

Using both DNS and large-eddy simulation (LES), Yang and Lei (1998) examined the role of turbulent scales in the settling velocity of heavy particles. They found that large-scale flow plays a significant role in determining the increase in settling velocity, namely, the increase in settling scales with rms fluid velocity (u'), not  $v_K$ , but peaks at  $St \approx 1$ .

The problem of settling rate of inertial particles has also been investigated experimentally. Zhou and Cheng (2009) monitored the settling velocity of low-inertia solid particles in turbulent water flow generated by an oscillating grid. They found that the actual settling velocity of the particles is generally smaller than the corresponding terminal velocity. The authors showed that the reduced settling velocity correlates with the vertical velocity fluctuation. Ruiz et al. (2004) measured the settling velocity of phytoplankton cells in turbulent flows generated in water by different mechanical devices. They concluded that the larger settling velocity observed in the experiment is likely due to mechanism of preferential sweeping. Cuthbertson and Ervine (2007) investigated the behavior of fine sand particles in turbulent open channel flow generated over rough, porous bed condition. Based on the experimental data they concluded that the settling rate of the suspended fine sand can be significantly enhanced over the terminal velocity.

Other mechanisms for enhanced settling, discussed in the literature, include the influence of local particle-particle hydrodynamics interactions (Aliseda et al., 2002; Alipchenkov and Zaichik, 2009) and the effect of two-way coupling (Dejoan, 2011; Yang and Shy, 2005; Bosse et al., 2006). There are also a number of turbulence-induced mechanisms that lead to reduction of particle settling velocity. The first concerns the effect of nonlinear drag. This mechanism has been investigated numerically in several previous studies. A first rigorous numerical simulation was developed by Wang and Maxey (1993). Their DNS were limited only to small turbulent Reynolds numbers and to three particle Stokes numbers. They found that the drag nonlinearity slightly reduces the net increase in the settling velocity.

In another study, Mei (1994) showed that for a nonfrozen turbulence the nonlinear drag increases the effective particle time constants and this leads to an effective reduction in the settling velocity. Similar conclusions result from the study by Fung (1993) in which a turbulent flow is modeled using kinematic simulations (KS). In a subsequent study, Chan and Fung (1999) investigated the gravitational settling of small inertial particles in a simple, twodimensional periodic cellular flow. They found that the settling velocity of inertial particles in such a flow is larger than the terminal velocity by 120% (assuming  $V_T/\sigma_u < 0.5$ ). It should be added that their particle equation of motion included the square drag. The effect of nonlinear drag on the particle settling velocity has also been examined by Stout et al. (1995). They demonstrated that reduction in the average settling velocity can be greater than 35%, and the rate of reduction decreases with decreasing particle Reynolds number. The main cause here is the decreasing drag nonlinearity and the reduction vanishes within the Stokes drag regime.

The results presented by Mei (1994), Fung (1993) and Stout et al. (1995) are in contrast to those developed by Wang and Maxey (1993) and Chan and Fung (1999). The discrepancy may be caused by the differences in the numerical approach employed to simulate the turbulent flows. In the literature, there are also some studies stating that the impact of the nonlinear drag on the settling velocity is negligible. For example, Nielsen (1993) derived an analytical formula that relates the maximum fluid acceleration to the reduction in the settling velocity. Based on this formulation and experimental measurements, Nielsen concluded that the effect of nonlinear drag on the settling velocity is less than  $10^{-5}V_T$  and hence has no practical importance.

The second mechanism for reduced settling is the vortex trapping. Tooby et al. (1977) showed experimentally that forced vortices with horizontal axes can trap bubbles and heavy particles and thus eliminate their settling velocity completely. Manton (1974) derived equations of motion for particles moving through an axisymmetric eddy with a horizontal axis. Based on this analysis, Manton claimed that the turbulence should cause particles to fall at much lower rate than their terminal velocity, even in the absence of a mean updraft. Nielsen (1993) proposed a basic theoretical model to explain the mechanism of vortex trapping. Vilela and Motter (2007) examined the trapping mechanism of aerosol particles in two idealized open flows such as the blinking vortex system and the leepfrog system. Their numerical simulations showed that for a broad range of conditions the permanent trapping of the inertial particles is possible even when the gravitational effect is large. Pasquero et al. (2003) investigated the settling of heavy particles in a steady two-dimensional random velocity field. Their analytical considerations show that particle suspension takes place when small-scale turbulence suitably alters the curvature of the streamlines around an eddy. They also argued that when the flow is unsteady, permanent suspension of particles becomes very unlikely. The above flows are perhaps not strictly relevant to 3D homogeneous isotropic turbulence, and therefore the latter results may not be directly generalized to 3D turbulence.

The third reduced-settling mechanism is known as the loitering effect. This mechanism was first observed and described by Nielsen (1993). Nielsen observed that in a relatively weak turbulence the settling velocity of heavy particles is reduced by 20– 40%. He argued that this delaying effect results from the fact that fast-falling particles spend more time in the regions with the upward flow. Such particles cannot be trapped by eddies but also cannot be effectively guided along vortical structures. This finding was confirmed in subsequent laboratory experiments by Kawanisi and Shiozaki (2008) and Cadiergue et al. (1999). Good et al. (2014) claimed that this effect is not observed in DNS simulations when the Stokes drag is assumed.

In this study, we employ direct numerical simulations to investigate the settling velocity of non-interacting small heavy particles. The first objective is to explore how the settling velocity varies with the non-dimensional parameters, namely *St*, *S<sub>V</sub>*, *Fr*, *R<sub>λ</sub>*,  $\sigma$  and turbulence intensity (i.e. u') or equivalently the energy dissipation rate ( $\varepsilon$ ) at fixed *R<sub>λ</sub>* and kinematic viscosity. To a limited extent, similar studies have been carried out in the past (see comparison in Table 2).

The second objective is to develop a parameterization (analytical formula) that relates the settling velocity to the set of nondimensional parameters indicated in Eq. 4. The third objective is to quantify how the large-scale forcing scheme, gravitational acceleration and the drag nonlinearity influence the settling statistics. Finally, the fourth objective is to explore the possible mechanisms by which the contributing parameters alter the particle sedimentation velocity.

The remainder of the paper is organized as follows. In Section 2, we describe the essential details of the numerical method. Results from the DNS are discussed in Section 3. These include sensitivity of the settling rate to *St*, *S*<sub>V</sub>, *Fr*,  $\sigma$ ,  $\varepsilon$ , u',  $R_{\lambda}$ , large-scale forcing scheme, gravitational acceleration, and the drag nonlinearity. Numerical experiments with artificially forced loitering are discussed in Section 4. Section 5 contains a summary and main conclusions.

## 2. Methodology

The turbulent fluid flow is simulated by a standard pseudospectral method (Orszag and Patterson, 1972), which solves the Navier-Stokes equation on a 3D uniform mesh with *N* equally spaced grid points in each spatial direction. The flow domain is a cube with size  $2\pi$ . Periodic boundary conditions are enforced which is consistent with the 3D discrete Fourier transform applied to the fluid velocity field.

The fluid velocity  $\mathbf{U}(\mathbf{x}, t)$  is obtained by numerical integration of the rotational form of the Navier–Stokes equation

$$\frac{\partial \mathbf{U}}{\partial t} = \mathbf{U} \times \boldsymbol{\omega} - \nabla \left(\frac{P}{\rho} + \frac{1}{2}\mathbf{U}^2\right) + \nu \nabla^2 \mathbf{U} + \mathbf{f}(\mathbf{x}, t).$$
(7)

for an incompressible fluid satisfying the continuity equation

$$\nabla \cdot \mathbf{U}(\mathbf{x},t) = \mathbf{0}.\tag{8}$$

Here  $\boldsymbol{\omega} \equiv \nabla \times \mathbf{U}$  is the vorticity, *P* is the pressure,  $\rho$  is fluid density and  $\nu$  is fluid kinematic viscosity.

A stationary turbulence is maintained by the forcing term  $\mathbf{f}(\mathbf{x}, t)$  which is nonzero only for a few low-wave-number modes ( $|k| < \sqrt{8}$ ) in the Fourier space. We implemented two forcing methods allowing us to study the sensitivity of the results on the forcing method (as discussed by Rosa et al. (2015)). The details of the two forcing methods are described in Section 3.4.

We first evolve the flow (without particles) from t = 0 up to at least  $t = 10T_e$  ( $T_e$  is the large-eddy turnover time) to ensure that the flow is statistically stationary. This eliminates any effect of initial flow condition on the statistics and dynamics of the resulting stationary homogeneous and isotropic turbulence. Detailed flow statistics obtained at different grid resolutions and different forcing mechanisms are presented in our preceding article (Rosa et al., 2013). Here, in Table 3, we only present a few basic parameters, namely, the kinematic fluid viscosity  $\nu$ , the energy dissipation

# Table 2

Overview of previous DNS studies investigating the settling velocity of small inertial particles in homogeneous isotropic turbulence. The table compares parameter setting and average flow statistics. E/R refers to the nature of obtained results, namely (E)nhancement or (R)eduction of the settling rate (B - both).

Study	$d_p/\eta$	St	Sv	Fr	$ ho_p    ho_f$	Coupling	$R_{\lambda}$ up to	Ν	Nonlin. drag	E/R/B
Wang and Maxey (1993)	0.01-0.2	0.2-3.25	0.2-3.25	0.008-34.33	1000	1-way	61.5	64	Yes	E
Yang and Lei (1998)	-	0.1-2.5	0.1-3.5	0.001-30.625	1000	1-way	65.3	96	Yes	E
Bosse et al. (2006)	-	$\sim 1$	$\sim 1$	$\sim 1$	5000	2-way	42.7	64	No	E
Ayala et al. (2008)	0.011-0.101	0.032-2.282	0.626-12.516	0.0124-357.48	1000	1-way	72.4	128	No	E
Dejoan (2011)	0.036-0.085	$\sim 1$	$\sim 1$	$\sim 1$	5000	2-way	130	512	No	E
Good et al. (2014)	0.012-0.55	0.03-60.0	0.1-100	0.0003-600,000	877	1-way	227	512	Yes	В
Present study	0.013-4.02	0.007-8.97	0.355-63.1	0.0002-12,618	500-10,000	1-way	500	1024	Yes	E

#### Table 3

Parameters and realized statistics (in spectral units) at the statistically stationary stage of the simulated flows using both deterministic and stochastic forcing schemes.

Ν	32	64	128	256	512	1024				
	Deterministic forcing scheme									
$\nu$ $\times$ $10^4$	170	67	28	11	4.8	1.8				
$\varepsilon \times 10^1$	1.52	1.79	2.08	2.00	1.91	1.96				
$\eta~ imes~10^2$	7.553	3.604	1.805	0.9033	0.4912	0.2339				
$\tau_K \times 10^1$	3.357	1.939	1.164	0.7420	0.5030	0.3040				
T <sub>e</sub>	3.811	3.848	3.636	3.774	3.946	3.917				
$u' \times 10^1$	7.58	8.28	8.67	8.68	8.65	8.73				
$k_{max}\eta$	1.095	1.099	1.128	1.143	1.250	1.194				
	Stochast	ic forcing sc	heme							
$\nu \times 10^3$	450	200	75	28	12	4.5				
ε	3334	3449	3504	3673	3670	3751				
$\eta   imes  10^2$	7.244	3.909	1.867	0.8810	0.4666	0.2223				
$\tau_K \times 10^2$	1.167	0.7646	0.4651	0.2774	0.1815	0.1099				
$T_e \times 10^2$	8.362	9.531	10.14	10.3	10.6	8.444				
u′	16.65	18.07	18.76	19.36	19.61	19.85				
$k_{max}\eta$	1.050	1.192	1.167	1.114	1.188	1.135				

Tab	le	4	
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Basic properties of cloud droplets.

	$\varepsilon = 10 \text{ cm}^2/\text{s}^3$		ε <b>=</b> 400	$\varepsilon = 400 \text{ cm}^2/\text{s}^3$		$\epsilon$ =1000 cm <sup>2</sup> /s <sup>3</sup>	
a [µm]	St	S <sub>V</sub>	St	S <sub>V</sub>	St	S <sub>V</sub>	Fr
10	0.010	1.12	0.063	0.45	0.100	0.355	0.013
15	0.023	2.52	0.143	1.00	0.226	0.798	0.144
20	0.040	4.49	0.254	1.78	0.401	1.419	0.808
22.5	0.051	5.68	0.321	2.26	0.507	1.796	1.64
27.5	0.076	8.48	0.479	3.37	0.758	2.683	5.46
30	0.090	10.1	0.571	4.01	0.902	3.193	9.20
40	0.160	17.9	1.014	7.14	1.604	5.676	51.7
50	0.251	28.0	1.585	11.1	2.506	8.869	197
60	0.361	40.4	2.283	16.0	3.609	12.77	589

rate  $\varepsilon$ , the Kolmogorov length  $\eta$ , the Kolmogorov time  $\tau_K$ , the eddy turnover time  $T_e$ , and the rms fluid velocity u'. Since the mean flow velocity in the domain is zero, u' defines also the turbulence intensity. The spatial resolution of the simulations was monitored by  $k_{max}\eta$ . In pseudo-spectral simulations this value should be greater than unity for fine scales to be resolved. In order to maximize turbulent Reynolds number  $(R_{\lambda})$  we kept  $k_{max}\eta$  close to unity for every mesh size and for both forcing methods (see Table 3).

The motion of heavy inertial particles is tracked by a Lagrangian approach. Once the background turbulent velocity field  $\mathbf{U}(\mathbf{X}, t)$  is established, inertial particles are advanced by solving their equation of motion (Maxey and Riley, 1983) including particle inertia, viscous drag, and the body force

$$\frac{d\mathbf{V}^{(k)}(t)}{dt} = -f(Re, V_{rel})\frac{\mathbf{V}^{(k)}(t) - \mathbf{U}(\mathbf{Y}^{(k)}(t), t)}{\tau_p^{(k)}} + \mathbf{g}$$
(9)

$$\frac{d\mathbf{Y}^{(k)}(t)}{dt} = \mathbf{V}^{(k)}(t).$$
(10)



**Fig. 1.** The increase  $(\langle V_S \rangle - V_T)$  in the particle mean settling velocity normalized by the terminal velocity as a function of the Stokes number. The simulations have been performed assuming  $V_T/v_K = 1$ . The results from present DNS simulations (blue color) are compared with results developed by Wang and Maxey (1993) and Bosse et al. (2006).

Here k is the particle number,  $\tau_p^{(k)}$  is the Stokes inertial response time,  $\mathbf{V}^{(k)}(t)$  is the particle velocity and  $\mathbf{U}(\mathbf{Y}^{(k)}(t), t)$  denotes the fluid velocity at the particle location  $\mathbf{Y}^{(k)}(t)$ , *Re* is particle Reynolds number,  $V_{rel}$  is particle-fluid relative velocity and **g** is the gravitational acceleration. *f* is the drag coefficient which in this paper is set to f = 1 except in Section 3.6 for which a non-linear drag is introduced and discussed. Particles are initially introduced to the flow at random locations. In this study we focus mainly on inertial particles relevant to cloud droplets of radius from 10 to 60  $\mu$ m.

A complete description of the code together with results of the numerical experiments and scalability analysis can be found in (Ayala et al., 2014). In this study, we consider only the one-way coupling, i.e. the fluid flow is not affected by the presence of particles.

Our DNS implementation is validated in two series of simulations performed at low Taylor microscale Reynolds number ( $R_{\lambda}$  = 27.8 and 48.4). The setup of the experiments follows the study by Wang and Maxey (1993). For this validation, it is assumed that the terminal velocity of the particles is equal to the Kolmogorov velocity ( $S_V = 1$ ). The focus is on computing the settling velocity of small rigid spherical particles with the Stokes number ranging from 0.2 to 4. In Fig. 1 we compare results from present DNS with results reported in Wang and Maxey (1993) and Bosse et al. (2006). The figure shows an increase in the particle mean settling velocity normalized by terminal velocity for different values of the Stokes number. We conclude that the present results are consistent with the results reported in Wang and Maxey (1993) and Bosse et al. (2006). The small differences at low grid resolution  $32^3$  are associated with the differences in  $R_{\lambda}$  and also the details of large-scale flow forcing.

# 3. The results of numerical simulations

3.1. Settling velocity of inertial particles relevant to cloud droplets in turbulent flow with different  $\varepsilon$ 

In the first series of numerical experiments, we assume that the particles are water droplets of density  $\rho_p = 1 \text{ g/cm}^3$ . The particles move in a turbulent air of density  $\rho_f = 0.001 \text{ g/cm}^3$  and kinematic viscosity  $\nu = 0.17 \text{ cm}^2/\text{s}$ . The vertical downward acceleration is set to  $\mathbf{g} = 9.8 \text{ m/s}^2$  and the average energy dissipation rate of the turbulent flow is in the range of 10 to 1000 cm<sup>2</sup>/s<sup>3</sup>. Such a setting is typical for the atmospheric cloud processes (Siebert et al., 2006). The setting can also be described in terms of three non-dimensional parameters ( $Fr = StS_V^2$  is not an independent parameter) specified in Table 1. The Taylor microscale Reynolds number is constant ( $R_{\lambda} = 143.74$ ) for each simulation. The values of these dimensionless parameters for three different energy dissipation rates are listed in Table 4.

Since both  $R_{\lambda}$  and the fluid viscosity  $\nu$  are fixed in the simulations, the relation between energy dissipation rate and turbulence intensity is uniquely defined by

$$u' = (\varepsilon \nu)^{0.25} \frac{\sqrt{R_{\lambda}}}{15^{0.25}} \propto \varepsilon^{1/4}.$$
(11)

Hence the dependence of the settling velocity on the dissipation rate can be equivalently expressed as dependence on the turbulence intensity (u').

Results from similar numerical simulations, but limited to low turbulent Reynolds numbers ( $R_{\lambda} \leq 72.4$ ) and to two energy dissipation rates (100 and 400 cm<sup>2</sup>/s<sup>3</sup>) have been reported in the study by Ayala et al. (2008). The present simulations were performed at a finer grid resolution of 256<sup>3</sup> and consequently, the Taylor microscale Reynolds number of the present DNS is approximately twice larger ( $R_{\lambda} = 143.74$ ) than those of Ayala et al. (2008). Additionally, the numerical uncertainty of the current results is smaller than those presented in Ayala et al. (2008). We run the DNS longer than 10 $T_e$  with up to one million particles. The long simulation time and the large number of particles resulted in the small statistical uncertainty.

The effect of turbulence on the settling rate is quantified by comparing the settling velocities of particles in turbulent flow against those in a stagnant air (i.e. terminal velocity). Fig. 2 shows the increase in the particle settling velocity normalized by the terminal velocity as a function of both particle (water droplet) radius and non-dimensional parameter Fr. Different lines correspond to simulations performed with different value of energy dissipation rate (equivalently u').

Several conclusions can be drawn from Fig. 2. First, the settling velocity of low inertia particles tends to zero, consistent with the tracer limit in which particles behave like fluid elements. We note that the mean flow velocity in the domain is zero. Second, the maximum increase in the settling velocity is observed at Froude number equal to 1 and is independent of the energy dissipation rate, consistent with the theory of Davila and Hunt (2001) and the results of Ayala et al. (2008). Third, the dynamics of droplets with large  $S_V$  ( $r > 60 \ \mu m, S_V > 10$ ) is weakly modified by turbulence due to very short interaction time with turbulent eddies (Fr > 500). Therefore, for these particles the settling velocity is comparable with their terminal velocity. Also in



**Fig. 2.** The increase in settling velocity of inertial particles in turbulent air normalized by their terminal velocity. Different colors correspond to simulations with different energy dissipation rate or equivalently turbulence intensity. For each series of simulations at fixed  $\varepsilon$ , the maximal increase of velocity occurs at  $Fr \approx 1$ . All simulations were performed at the grid resolution of 256<sup>3</sup> ( $R_{\lambda} = 143.74$ ) using the stochastic forcing.

Fig. 2, the empirical formula of Ghosh and Jonas (2001) is plotted for comparison. Their model predicts the amplification effect as  $(\langle V_S \rangle - V_T)/V_T = \alpha^{-\beta a}$ , where  $\alpha = 5.5$  and  $\beta = 0.173 \ \mu m^{-1}$ . The differences between DNS and the empirical formula are large. Even the trend for large inertia particles is only roughly predicted. The empirical formula was developed by curve-fitting an exponential function to a few discrete data points computed based on the theoretical model. Therefore, this function cannot correctly represent amplification effect both for small and large droplets.

The DNS results from Fig. 2 can be presented equivalently as a function of two other nondimensional parameters, namely *St* and *S<sub>V</sub>*. The contour plot in Fig. 3 shows the increase in the settling velocity normalized by u' as a function of both  $\ln St$  and  $\ln S_V$ . The black circles represent the data points from various simulations and the two-dimensional distribution is obtained by spatial interpolation. Compared to the previous figure, Fig. 3 shows also results from two additional series of simulation performed at  $\varepsilon = 5$  and 30 cm<sup>2</sup>/s<sup>3</sup>.

Only in two cases, we observed a reduction of the settling velocity. For particles with low inertia of St = 0.007 and small settling of  $S_V = 1.33$  the reduction is about 0.18 mm/s while for particles with St = 0.143 and larger  $S_V = 27.01$  the observed reduction is 4 mm/s. These values are relatively small and are comparable in magnitude with the statistical uncertainty. We note that the results from present DNS in the range St > 1.5 differ from those obtained experimentally by Yang and Shy (2003), namely, reduction of the settling velocity is not observed in DNS data. Yang and Shy measured the settling velocity of heavy tungsten and glass particles in an aqueous near-isotropic turbulence. They stated, that the settling velocity of particles with St > 1.5 is smaller than their terminal velocity. This reduction results from the nonlinear drag effect due to both larger particle Reynolds number and interaction with larger turbulent scales. Yang and Shy (2003) observed the reduction in  $\langle V_{\rm S} \rangle$  only for glass particles with a large particle Reynolds number



**Fig. 3.** Difference between the average settling velocity of inertial particles relevant to cloud droplets in turbulent flow and its terminal velocity in stagnant air normalized by u'. This difference is presented as a function of  $\ln(St)$  and  $\ln(S_V)$ .

 $Re = (2aV_T/v) = 39$ . In our DNS, however, the maximal Re is significantly lower at Re = 15.07.

A similar conclusion is reported in a recent study by Good et al. (2014). Good et al. (2014) performed experiments with water droplets in turbulent air in which the turbulence was generated using 32 loudspeaker jets. Their experiments show that for certain particle classes there is a reduction of the settling velocity. They also concluded that such a reduction is not observed in DNS results when the particles are subjected to the linear drag.

All DNS discussed in this section were performed at the same grid resolution (N = 256). Thus, both  $R_{\lambda}$  and  $u'/v_K$  are fixed for each simulation. Therefore, the changes in the energy dissipation rate can be viewed equivalently as changes in the particle inertia. To confirm this interdependence, the original data from Fig. 2 needs to be presented in another nondimensional form. Fig. 4 shows the relative increase of the settling velocity of the particles normalized by the Stokes number versus the particle Froude number. With such scaling all the lines collapse to one curve. This normalization turns out to be helpful in developing a useful parameterization. Based on Eq. 4, we postulate that the increase in the settling velocity takes on the following form

$$\frac{\langle V_S \rangle - V_T}{V_T} = \frac{\mathbf{F}_1(St, S_V, R_\lambda, ...)}{S_V} = St \cdot \mathbf{F}_4(Fr) \cdot Fr^{m(R_\lambda)} \cdot \mathbf{F}_5(R_\lambda) \quad (12)$$

where the product of  $Fr^{m(R_{\lambda})}$  and  $\mathbf{F}_5(R_{\lambda})$  is equal to unity at  $R_{\lambda} =$  143.74. In other words, we decompose the function  $\mathbf{F}_1(St, Sv, R_{\lambda})$  from Eq. 4 into two parts (i.e.  $\mathbf{F}_4$  and  $\mathbf{F}_5$ ), the first of which depends only on *Fr*, and the second one includes effects of  $R_{\lambda}$ . An important advantage of this decomposition is that the effects of inertia are separated from the effects of the flow Reynolds number. For the simulations discussed in this section, Eq. 12 reduces to

$$\langle V_{\rm S} \rangle - V_{\rm T} = V_{\rm T} \cdot St \cdot \mathbf{F}_4(Fr). \tag{13}$$

By curve fitting we obtain the analytical formula of  $F_4$ , namely

$$\mathbf{F}_4(x) = \frac{p_1 x^2 + p_2 x + p_3}{x^2 + q_1 x + q_2} \tag{14}$$



**Fig. 4.** The difference between the average settling velocity and terminal velocity normalized by terminal velocity and Stokes number as a function of particle Froude number.



**Fig. 5.** The difference between the average settling velocity and terminal velocity normalized by terminal velocity as a function of particle Froude number. Markers indicate results obtained from DNS. Dashed lines (with matching colors) refer to the new parameterization.

where  $x = \log_{10}(Fr)$ ,  $p_1 = 0.8296$ ,  $p_2 = -4.184$ ,  $p_3 = 5.316$ ,  $q_1 = 0.2412$  and  $q_2 = 5.897$ .

The accuracy of the new parameterization is verified by comparing against DNS data. The comparison is presented in Fig. 5. We find a good quantitative agreement between the parameterization and DNS results for a wide range of the particle Froude number. In Section 3.3 we will extend this parameterization to different  $R_1$ .

Dissipation of energy represents in part the intensity of local mixing by large scales in the turbulent flow. In isotropic homogeneous turbulence the average energy dissipation rate can be approximated using the large scale characteristics (Mouri et al., 2012), namely

$$\langle \varepsilon \rangle \sim \frac{u'^3}{L_f}$$
 (15)

#### Table 5

Turbulence statistics (in physical units) of the simulated flows at 256<sup>3</sup> using the stochastic forcing scheme and normalized droplet properties. The response time  $\tau_p$  and terminal velocity  $V_T$  correspond to droplets with a radius of 20  $\mu$ m.

$\varepsilon$ [cm <sup>2</sup> /s3]	<i>u'</i> [cm/s]	$L_f$ [cm]	$T_e$ [s]	$\tau_p/T_e$	$V_T / u' * 10^2$
		,		1	
10	6.96	33.58	4.84	0.74	0.11
50	10.41	22.46	2.17	0.49	0.24
100	12.38	18.88	1.53	0.41	0.34
200	14.72	15.88	1.08	0.35	0.48
400	17.51	13.35	0.77	0.29	0.68
600	19.37	12.07	0.63	0.26	0.83
800	20.82	11.23	0.54	0.25	0.97
1000	22.01	10.62	0.48	0.23	1.09



**Fig. 6.** The difference between the average settling velocity and terminal velocity normalized by terminal velocity as a function of square root of energy dissipation rate and  $u'^2$ . Different colors and markers correspond to different size of the inertial particles (cloud droplets).

where  $L_f$  represents the size of the energy-containing eddies and can be set to the longitudinal integral length scale. Table 5 shows the mutual relation between the energy dissipation rate, the rms fluctuating velocity and  $L_f$ . The increase of  $\varepsilon$  form 10 to 1000 cm<sup>2</sup>/s<sup>3</sup> results in a threefold decrease of  $L_f$  and a threefold increase of u'. The increase of u' is the main factor affecting the settling velocity of inertial particles. Wang and Maxey (1993) showed that particles settling in the channel-like regions are indeed strongly affected by the velocity at large turbulent scales such as u'.

In this study we found that for a wide range of particle Froude numbers the relation between  $u'^2$  and  $(\langle V_S \rangle - V_T)$  is linear (see Fig. 6). Since in all discussed DNS  $u' \propto \varepsilon^{1/4}$  the increase in the settling rate is also a linear function of  $\sqrt{\varepsilon}$ .

Interestingly, by changing the flow dissipation rate the relative importance of gravity vs particle inertia is modified. That is because both *St* and *S*<sub>V</sub> are functions of  $\varepsilon$ , namely  $St \propto \varepsilon^{1/2}$  and  $S_V \propto \varepsilon^{-1/4}$ . However, it does not imply that  $(\langle V_S \rangle - V_T)$  is also a linear function of *St*. The reason is that the Stokes number depends also on the particle size, namely  $St \propto a^2$ .

An important question to pose here is whether the linear dependence on  $\sqrt{\varepsilon}$  is general or a characteristic only for this partic-



**Fig. 7.** The difference between the average settling velocity and terminal velocity normalized by terminal velocity as a function of square root of energy dissipation rate. Simulations have been performed at four different  $R_{\lambda}$  and droplets of size  $a = 20 \ \mu m$  (solid lines) and  $a = 35 \ \mu m$  (dashed lines).

ular  $R_{\lambda}$ . To investigate the effect of Reynolds number, additional series of simulations at different  $R_{\lambda}$  and different particle *Fr* have been performed. Results from these DNS are presented in Fig. 7. We confirm that linearity of the relationship does not depend on  $R_{\lambda}$ . However, the increase of the settling rate is larger at higher  $R_{\lambda}$ . This issue will be analyzed more broadly in Section 3.3.

# 3.2. The effect of particle-to-fluid density ratio

In the previous section we considered conditions which are characteristic for cloud processes in the atmosphere. In these simulations, we kept the density ratio fixed and equal to  $\sigma = 1000$ . The results do not show a clear evidence of a reduction in the settling rate of small cloud droplets. In this section, we discuss the results of DNS performed at different density ratio in the range of 500 <  $\sigma$  < 10000. To perform simulations at even lower density ratio (i.e.  $\sigma$  < 500), other hydrodynamic forces (e.g., history term, added mass, etc.) will have to be included. The new set of simulations was performed at fixed energy dissipation rate  $\varepsilon = 800 \text{ cm}^2/\text{s}^3$ . Other parameters such as gravity, fluid viscosity and density remain unchanged namely,  $\mathbf{g} = 9.8 \text{ m/s}^2$ ,  $\nu = 0.17 \text{ cm}^2/\text{s}$  and  $\rho_f = 0.001 \text{ g/cm}^3$ .

Fig. 8 shows the increase  $(\langle V_S \rangle - V_T)$  in the particle mean settling velocity normalized by the terminal velocity. Different colors correspond to different density ratio (with blue curve representing water droplets in turbulent air). Several conclusions emerge from Fig. 8. First, regardless of the density ratio, the average settling velocity is always larger or equal to the terminal velocity. Second, the maximum increase in the settling velocity does not exceed 36%. This is a consequence of a finite and fixed value of the energy dissipation rate that is used in all simulations. Third, particles of high density (large  $\sigma$ ) are less responsive to the turbulence eddies and therefore an increase in the settling velocity occurs only for particles with small radii (inertia).

The shape of the functions corresponding to different  $\sigma$  is generally preserved. It is worth mentioning that the particle equation of motion can be rewritten in dimensionless form in terms of *St* and *S<sub>V</sub>*. In our simulation the ratio  $St/S_V = \varepsilon^{3/4}/(gv^{1/4})$  is constant and equal to 0.24. Therefore, the results will collapse to a single



**Fig. 8.** Increase in the particle mean settling velocity normalized by the terminal velocity as a function of particle radius. Different markers (colors) correspond to different density ratio ( $\sigma = 1000$  - water droplets in air). Each simulation was performed assuming a fixed energy dissipation rate  $\varepsilon = 800 \text{ cm}^2/\text{s}^3$ . Resolution of computational grid was 256<sup>3</sup>.

curve when plotted as a nondimensional parameters St and  $S_V$  (not shown for conciseness).

#### 3.3. The effect of flow Taylor microscale Reynolds number

Another important parameter that affects the settling velocity of inertial particles is the range of turbulent scales present in the flow. Typically, the range of scales is defined by the Taylor microscale Reynolds number  $(R_{\lambda} = u'\lambda/\nu)$ . In pseudo-spectral simulations  $R_{\lambda}$  depends, among other things, on resolution of the computational grid, dealiasing truncation, forcing scheme and fluid viscosity. The DNS results discussed in previous section were limited to only one grid resolution (256<sup>3</sup>) and the Taylor microscale Reynolds number was fixed ( $R_{\lambda} = 143.74$ ). To examine the effect of  $R_{\lambda}$  on the settling rate, more simulations at various grid resolutions is needed. Performing DNS at large computational grids (e.g. 1024<sup>3</sup>) is numerically expensive and requires advanced and efficient codes adapted to modern supercomputing architectures (Ayala et al., 2014). The high computational cost was a major obstacle to address this problem in the past.

Our present DNS cover a wide range of flow Taylor microscale Reynolds numbers (up to 373 in simulations with stochastic forcing and up to 500 in simulations with deterministic forcing). The new simulations were performed at various grid resolutions ( $32^3$ –  $1024^3$ ). First, we analyze the settling velocity of low inertia particles relevant to small water droplets with 5 to 20  $\mu$ m radii. Other basic parameters are as follows:  $\mathbf{g} = 9.8 \text{ m/s}^2$ ,  $\sigma = 1000$  and  $\nu = 0.17 \text{ cm}^2/\text{s}$ .

The results obtained in DNS are presented in Fig. 9a with markers. The increase of the settling velocity normalized by  $V_T$  is plotted as a function of both  $R_{\lambda}$  and grid size (*N*). The energy dissipation rate is set to  $\varepsilon = 400 \text{ cm}^2/\text{s}^3$ . The results indicate that the increase in the settling rate of low inertia particles ( $a = 10 \mu \text{m}$ , Fr = 0.013) is insensitive to  $R_{\lambda}$ . An increase in the Reynolds number from 27.8 to 225 results only in 5% increase in the normalized settling velocity. The main observation concerns the saturation at large  $R_{\lambda}$ . This suggests that the particle response time is



**Fig. 9.** The increase in average settling velocity normalized by  $V_T$  as a function of Taylor microscale Reynolds number and size of the computational grid. Markers indicate results obtained from DNS. Dashed lines (with matching colors) refer to the parameterization. (a) Results for particle radii in the range 10–20  $\mu$ m. The increase in settling velocity grows with *Fr*. (b) Results for particle radii in the range 20–60  $\mu$ m. For this range, unlike (a), the increase in settling velocity decreases with *Fr*.

much smaller than the characteristic time scale of the largest eddies. For small  $R_{\lambda}$  the settling velocity converges to the terminal velocity because the kinetic energy of small turbulent structures is not sufficient at this  $R_{\lambda}$  to alter the particle settling velocity. For the particles of size  $10 < a < 20 \ \mu m$  (0.013 < Fr < 0.808) a saturation is still observed at a higher  $R_{\lambda}$ . Additionally the increase in terminal velocity is larger, which is consistent with the results presented in Fig. 2.

Our results for larger particles (25-60  $\mu$ m in radii, 1.6 < *Fr* < 589) are shown in Fig. 9b. There are two essential differences between settling velocity of small and large particles. Regardless of the Reynolds number, the actual settling velocity tends to terminal velocity when the particle size increases. This is consistent with the conclusions from Fig. 2 and the results from large particle inertia *St* and large *S<sub>V</sub>*.

By increasing the grid resolution, or equivalently the domain size (for a given value of  $\varepsilon$ ), a larger turbulent scales are represented in the flow (see Table 6). This in turn leads to an increase in the settling rate of the particles with larger inertia. For example, the increase in the settling velocity of 60 µm particles is noticeable only when  $R_{\lambda} > 150$ . Yet, the increase is very small and does not exceed 2.4% at  $R_{\lambda} = 375$ .

#### Table 6

Basic characteristics of particles with respect to turbulent scales. Statistics are computed based on DNS data. The flow was forced by stochastic scheme. The energy dissipation rate was set to  $\varepsilon$ =400 cm<sup>2</sup>/s<sup>3</sup>.

a [µm]	25	30	40	50	60
St S <sub>V</sub>	0.40 2.79	0.57 4.01	1.01 7.14	1.58 11.15	2.28 16.06
$ au_p/T_e$ $V_T/u'$	$R_{\lambda} = 27.$ 0.055 1.04	.7, N = 32 0.080 1.50	2 0.142 2.66	0.221 4.16	0.319 5.99
$ au_p/T_e$ $V_T/u'$	$R_{\lambda} = 84.$ 0.017 0.59	.4, N = 12 0.025 0.86	28 0.045 1.52	0.070 2.39	0.101 3.43
$ au_p/T_e$ $V_T/u'$	$R_{\lambda} = 373$ 0.004 0.28	$\begin{array}{l} 3, \ N = \ 10 \\ 0.006 \\ 0.41 \end{array}$	24 0.010 0.72	0.016 1.13	0.023 1.64

The DNS data are helpful to complete the parameterization proposed in Section 3.1. The present results show that the increase in the settling rate depends on  $R_{\lambda}$  but this relationship is different for different Froude numbers. To separate the contribution of *Fr* (to  $\langle V_S \rangle$ ) which results from different Reynolds numbers we need to find a function  $m(R_{\lambda})$  for which all data points in Fig. 9 (normalized by *St* and **F**<sub>4</sub>) will collapse to a single curve. Given m = 0 at  $R_{\lambda} = 143.74$ , we propose the following formula for the exponent *m* 

$$m(R_{\lambda}) = c_1 \log_{10} \left(\frac{R_{\lambda}}{143.7}\right) + c_2 \log_{10} \left(\frac{R_{\lambda}}{143.7}\right)^2.$$
 (16)

The coefficients  $c_1 = 0.2125$  and  $c_2 = 0.0175$  are determined using 2D least square optimization.

The last step is to determine  $\mathbf{F}_5$ . We propose an analytical formula for  $\mathbf{F}_5$  by fitting the second order polynomial to the discrete DNS data. The fitted function has the following form

$$\mathbf{F}_{5}(\log_{10}(R_{\lambda})) = -0.4005(\log_{10}(R_{\lambda}))^{2} + 2.4285\log_{10}(R_{\lambda}) - 2.338$$
(17)

To verify the accuracy of the new parameterization we made a strict comparison of its values with the DNS data. In Fig. 9 the analytical predictions are plotted with dashed lines. The colors are consistent with the colors of markers used for plotting the DNS results. We confirm a good quantitative agreement between DNS and the parameterization for particles with low inertia (Fr < 0.8). The analytical formula predicts a small reduction of  $\langle V_S \rangle$  at large  $R_{\lambda}$ . For particles with medium Fr (i.e. 1 < Fr < 20) the parameterization yields a rather qualitative agreement. The trend at low  $R_{\lambda}$  is reproduced but the saturation at large  $R_{\lambda}$  is not correctly predicted. For particles with the Froude numbers 23.2 < Fr < 51.7 there is no saturation at large  $R_{\lambda}$  and again we observe a good quantitative agreement between DNS and the parameterization. For particles with large Froude numbers (Fr > 100) the increase in settling rate is very small (comparing to  $V_T$ ). Such small differences are qualitatively reproduced by the parameterization.

The results presented in Fig. 9 are useful in designing future numerical simulations. For instance one can estimate the optimal size of the computational domain for a given particle size. Modeling of turbulence at resolution 1024<sup>3</sup> or higher is numerically very expensive and in view of these results such effort may not be necessary.

#### 3.4. The effect of the large-scale forcing scheme

Turbulent flows are dissipative, therefore to maintain their stationarity a continuous supply of energy is required. In direct numerical simulations the turbulent flow is driven by an external

#### Table 7

The average settling velocity of cloud droplets in turbulent flow. The simulations were performed at six different grid resolutions using stochastic forcing scheme. Numerical uncertainties do not exceed 0.1%.

	$egin{array}{ll} N = \ R_{\lambda} = \end{array}$	32 28	64 48	128 84	256 144	512 225	1024 372			
а	$V_T [cm/s]$		$\langle V_S \rangle$ [cm/s]							
10	1.281	1.354	1.382	1.420	1.430	1.431	1.411			
12.5	2.001	2.182	2.245	2.322	2.343	2.332	2.322			
15	2.882	3.192	3.331	3.461	3.513	3.530	3.486			
17.5	3.923	4.414	4.612	4.829	4.928	4.918	4.890			
20	5.124	5.746	6.015	6.343	6.479	6.528	6.508			
22.5	6.485	7.185	7.548	7.959	8.198	8.247	8.207			
25	8.006	8.706	9.118	9.637	9.930	10.041	10.081			
27.5	9.688	10.318	10.785	11.372	11.806	11.919	11.995			
30	11.529	12.035	12.513	13.137	13.566	13.892	13.986			
35	15.693	15.959	16.278	16.896	17.499	17.959	18.252			
40	20.497	20.636	20.765	21.179	21.874	22.485	23.140			
45	25.941	26.061	26.031	26.270	26.773	27.498	28.324			
50	32.026	32.086	32.132	32.157	32.414	33.085	33.974			
60	46.118	46.137	46.180	46.184	46.207	46.521	47.197			

forcing to avoid turbulence decay. In pseudo-spectral methods the kinetic energy is injected in the Fourier space to low-wavenumber modes. Restricting the forcing to large scales only is thought to ensure that the small scales of the flow are unaffected by the forcing (Ishihara et al., 2009).

The simulations analyzed in previous sections have been performed using the stochastic forcing (Eswaran and Pope, 1988; Rosa et al., 2015) which is based on six Uhlenbeck-Ornstein random processes (Eswaran and Pope, 1988). The energy is continuously injected to the 80 low-wavenumber modes, namely ( $|\mathbf{k}| < \sqrt{8}$ ).

Since the forcing scheme affects the large scales (energetic scales) of the flow, it raises a question about sensitivity of the DNS results to the forcing method. In this section we address this question by comparing two sets of simulations performed with two different forcing methods. The supplementary set of simulations was performed with deterministic forcing (Sullivan et al., 1994). In deterministic forcing the energy levels of the two lowest wavenumber shells ( $0.5 < |\mathbf{k}| < 1.5$  and  $1.5 < |\mathbf{k}| < 2.5$ ) are fixed to E(1) = 0.55544 and E(2) = 0.159843, respectively.

Fig. 10 shows the average settling velocity of inertial particles in a turbulent flow normalized by the terminal velocity as a function of  $\ln(Fr)$ , particle radius,  $\ln(R_{\lambda})$  and normalized turbulent intensity. The simulations were performed assuming  $\mathbf{g} = 9.8 \text{ m/s}^2$ ,  $\nu = 0.17$  cm<sup>2</sup>/s,  $\sigma = 1000$  and  $\varepsilon = 400$  cm<sup>2</sup>/s<sup>3</sup>. Fig. 10 compares the results obtained with two different forcing schemes namely, (a) stochastic and (b) deterministic. Both methods show a similar trend, although there is a significant difference in the velocity magnitude for a given particle size and mesh resolution (N). At fixed resolution, particles settle faster in the flow forced by the deterministic scheme. This partially results from the different flow Reynolds numbers. Namely, the flow statistics show that the deterministic scheme leads to a higher flow Reynolds number. This in turn can be related to larger size of the energetic eddies which are responsible for particle acceleration. It should be noted that the results are consistent with the results presented in Fig. 9, where larger  $R_{\lambda}$  leads to larger  $\langle V_S \rangle$ .

To facilitate a quantitative comparison of the results obtained with different forcing schemes, the data presented in Fig. 10 are tabulated in Table 7 and 8.

## 3.5. The effect of gravity

Next, we extend the results beyond cloud droplets, allowing particle Stokes number and dimensionless sedimentation velocity  $S_V$  to vary independently. We note that the effects of gravity on



**Fig. 10.** The average settling velocity of inertial particles relevant to cloud droplets in turbulent air normalized by the terminal velocity as a function of  $\ln(Fr)$ , particle radius,  $\ln(R_{\lambda})$  and normalized turbulent intensity. DNS results obtained with (a) stochastic forcing, (b) deterministic forcing.

Tal	ble	8					
As	in	Table	7	but	for	deterministic	forcing.

	N =	32	64	128	256	512	1024			
	$R_{\lambda} =$	44	77	121	197	303	499			
а	$V_T [cm/s]$		$\langle V_S \rangle$ [cm/s]							
10	1.281	1.420	1.435	1.433	1.440	1.453	1.423			
12.5	2.001	2.283	2.353	2.355	2.355	2.386	2.399			
15	2.882	3.470	3.544	3.564	3.581	3.611	3.540			
17.5	3.923	4.768	4.949	4.970	5.005	5.079	5.075			
20	5.124	6.303	6.490	6.674	6.695	6.684	6.677			
22.5	6.485	7.941	8.182	8.355	8.423	8.501	8.501			
25	8.006	9.706	9.949	10.238	10.366	10.373	10.417			
27.5	9.688	11.468	11.827	12.104	12.237	12.417	12.496			
30	11.529	13.296	13.802	14.118	14.248	14.481	14.745			
35	15.693	17.435	17.901	18.361	18.545	18.880	19.311			
40	20.497	22.023	22.423	22.851	23.567	23.705	24.097			
45	25.941	27.057	27.533	28.003	28.901	29.244	29.375			
50	32.026	32.665	33.757	34.200	34.774	34.812	35.663			
60	46.118	45.922	47.001	47.675	48.477	48.907	49.852			



**Fig. 11.** The ratio of the average settling velocity of particles to the terminal velocity as a function of  $S_V$ . The particles at each set have a fixed Stokes number as in the legend. Different colors correspond to different *St*. The parameter  $S_V$  was adjusted by varying the gravitational acceleration. The simulations were performed at grid resolution of 256<sup>3</sup> assuming  $\varepsilon = 800 \text{ cm}^2/\text{s}^3$ .

particle acceleration and relative velocity in the independent St- $S_V$  space is recently studied in Parishani et al. (2015). Hence, here we extend and explore the results for different values of gravity to allow a varying  $S_V$  while keeping the Stokes number constant. Eight series of simulations were performed for a fixed set of parameters, namely, u' = 20.82 cm/s,  $\varepsilon = 800 \text{ cm}^2/\text{s}^3$ ,  $\nu = 0.17 \text{ cm}^2/\text{s}$  and  $\rho_p/\rho_f = 1000$ . In each series the Stokes number is determined by particle size. The range of considered Stokes numbers is 0.0224–1.815.  $S_V$  is a function of gravity and depending on particle inertia varies from 0.094 to 45.58. We assumed that the particle density corresponds to that of water ( $\rho_p = 1 \text{ g/cm}^3$ ) while the fluid density corresponds to that of air ( $\rho_f = 0.001 \text{ g/cm}^3$ ).

There are two limiting cases for this problem. First, for  $S_V = 0$  (corresponding to  $\mathbf{g} = 0$ ) and the second when  $S_V$  is large (equivalently  $\mathbf{g} \to \infty$ ). In both cases, the particle velocity should be equal to the terminal velocity. Specifically, when  $\mathbf{g} = 0$ , there is no featured direction and therefore both average particle velocity and terminal velocity are zero. For a large  $S_V$ , turbulence does not play a significant role and therefore the particle settling velocity should be equal to the terminal velocity. Nontrivial effect of gravity on the settling velocity of particles can be observed for intermediate values of  $S_t$  and  $S_V$ .

Fig. 11 shows the ratio of the average settling velocity of particles to the terminal velocity as a function of  $S_V$ . Each curve corresponds to a series of simulations with a fixed Stokes number. The settling velocity of particles with higher inertia St > 0.2 quickly converges to the terminal velocity (and hence a large slope) when gravity increases. This is the consequence of reduction in particle-eddy interaction time as particles settle faster at higher **g**. For low inertia particles (St = 0.0223 and 0.0504), however, we observe a slight increase in the settling velocity when gravity increases from 1**g** to 2**g** (the first two data points of each set). This increase may be due to larger Fr number ( $Fr \propto g^2$ ). On the other hand, in the limit  $S_V \rightarrow 0$ , low inertia particles behave like fluid elements and consequently  $\langle V_S \rangle \rightarrow V_T$ . For larger values of  $S_V$  the actual settling velocity gradually converges to the terminal velocity.

Table 9

Different formulation of drag factor.

Stokes dragf = 1Nonlinear drag (Clift et al., 1978) $f = 1 + 0.15Re^{0.687}$ Nonlinear drag (Stout et al., 1995) $f = \frac{C_0}{24}Re$  where  $C_D = \frac{1}{4} + \frac{24}{Re} + \frac{6}{\sqrt{Re}}$ Square drag (Fung, 1993) $f = \frac{|U-V|}{V_T} = Re\sqrt{\frac{\sigma}{18F_T}}$ 



**Fig. 12.** Drag coefficients as a function of particle Reynolds number  $Re = |\mathbf{U} - \mathbf{V}|d/\nu$ . Different colors correspond to different formulations. For square drag we assumed  $\sigma = 1000$ .

#### 3.6. The effect of drag nonlinearity

In the introduction we presented different and sometimes contradictory conclusions resulting from the previous studies in which a nonlinear drag was considered. Therefore those results may not be reconciled quantitatively as for the effect of nonlinear drag on the settling velocity. In the present study we address this problem by comparing the different formulations of the drag coefficients and number of simulations with different turbulence intensity and particle inertia (size).

The most common representations of the drag coefficient are: the Stokes drag, the standard nonlinear drag (Clift et al., 1978; Stout et al., 1995) and the square drag (Fung, 1993). We should point out that the square drag is unphysical for small particle Reynolds numbers. The mathematical formulas of these coefficients are given in Table 9. The two representations of the nonlinear drag show that f is a function of the particle Reynolds number only. The square drag depends additionally on the density ratio and the Froude number. For a more extended discussion of the drag factor, including turbulence effects, compressibility and different flow regimes we refer to ((Crowe et al., 2011) Chapter 4.3.1).

Fig. 12 shows a quantitative comparison of the drag coefficients computed using four different formulations as a function of the particle Reynolds number, density ratio and *Fr*. Interestingly, for the considered range of *Re*, there is only a slight difference between the Stokes drag and those of Clift et al. (1978) and Stout et al. (1995). However, the difference between the square drag and the three other formulations can be large, especially for small *Fr*. To predict the impact of the square drag on the settling velocity we should consider two extreme cases. First, when the particle has large inertia the drag coefficient goes to zero, so the settling velocity should be comparable with terminal velocity. Second, when particle has very low inertia the drag force acting is very large.

This makes the equation of motion unstable, and it is necessary to use a very short time step for integration of the equation of motion. After a relaxation period the droplets will behave as a fluid elements due to the strong drag force.

Fig. 13 shows the mean particle settling velocity computed in simulations with both Stokes drag and nonlinear drag (Clift et al., 1978). All simulations have been performed at 256<sup>3</sup> grid resolution with turbulent Reynolds number of  $R_{\lambda} = 143.74$ . Our results are in qualitative agreement with those of Wang and Maxey (1993) and we conclude that the differences in the average settling velocity computed with the different drag forces are relatively small. The differences do not grow with the turbulence intensity. We confirm that the nonlinear drag (Clift et al., 1978) only slightly reduces the net increase in the settling velocity.

The next step is to quantify the impact of the square drag on the settling rate. The large differences between drag coefficients presented in Fig. 12 may suggest that the effect of the square drag on the settling rate can be greater than those presented in Fig. 13. In Fig. 14 we compare the mean particle settling velocity (normalized by terminal velocity characteristic for the Stokes drag) computed using three different formulation of the drag force. All series of simulations have been performed at  $R_{\lambda} = 143.74$  and energy dissipation rate of  $\varepsilon = 400 \text{ cm}^2/\text{s}^3$ . We conclude that the difference in the settling rate computed using three different formulations of the drag force is small and does not exceed 1%. The most visible difference is observed between the Stokes drag and the square drag using particle (water droplets) of  $a \approx 20 \text{ }\mu\text{m}$  in radii.

The presented results do not show a significant reduction in the settling velocity if the particles are subjected to the nonlinear drag. For large particles (60  $\mu$ m) the visible differences (see Fig. 13) between simulations performed with Stokes and nonlinear drags are of order of the statistical uncertainty. In this regards, our results are different from those presented by Good et al. (2014). They showed that when the nonlinear drag is considered, the settling velocity of heavy inertial particles  $(S_{V_l} = \tau_p g/u' > 1)$ is smaller than the corresponding terminal velocity. The reduction is up to  $(V - V_T)/u' \approx -0.1$ . Interestingly, they obtained a similar trend both in DNS and in experiment. It is worth noting that the range of particle Reynolds number covered in their DNS (Fig. 3a therein) is approximately two times wider than in the present study. The main difference between the present work and Good et al. (2014) simulations lies in the flow forcing method. All our simulations with the nonlinear drag were performed using the stochastic forcing (Eswaran and Pope, 1988; Rosa et al., 2015) while Good et al. (2014) performed simulations using a deterministic forcing. In order to make a strict comparison with Good et al. (2014) study, we perform an additional set of simulations with the deterministic forcing scheme. Results from the later simulations are shown in Fig 15. Three series of simulations correspond to three values of energy dissipation rate, namely 100, 400 and 800 cm<sup>2</sup>/s<sup>3</sup>. Additionally we considered a wider range of particle radii i.e. 10-100 µm.

DNS results computed using deterministic forcing are consistent with those obtained with the stochastic forcing. We do not observe a clear evidence of reduction in the settling velocity of large inertial particles when the nonlinear drag is used. It is worth noting that we run the simulations at least  $100T_e$ . The statistics were collected after  $40T_e$  and averaged over at least  $60T_e$ . Good et al. (2014) averaged their statistics over about  $10T_e$ . The initial flow stabilization time and also the total time over which the averaging is performed may play a critical role in understanding the difference between our results and those presented by Good et al. (2014). Rosa et al. (2015) showed that the settling mechanism depends on particle inertia and  $R_{\lambda}$ . Large-inertia particles accumulate in the downward flow regions forming elongated (filamentlike) structures. The elongated structures are stable and remain for



**Fig. 13.** The mean settling velocity normalized by terminal velocity (computed with the Stokes drag) as a function of droplet radii. Different colors correspond to different turbulence intensity. Solid lines represent results computed with the Stokes drag. Results computed with nonlinear drag (Clift et al., 1978) are shown by dashed lines. The simulations were performed at grid resolution of 256<sup>3</sup>.



**Fig. 14.** Effect of the drag force on the settling rate. Different colors correspond to simulations with different drag coefficients.



**Fig. 15.** As in Fig 13 but for deterministic forcing. All series of simulations have been performed at the same grid size i.e.  $256^3$  and turbulent Reynolds number  $R_{\lambda} = 196.87$ .

a long time because the gravitational settling dominates the mechanism of turbulent mixing. We anticipate that for such elongated structures the different types of drag force, namely the nonlinear versus Stokes drag, may not have a significant impact on the settling velocity.

# 4. Effect of forced loitering

The DNS results presented in the previous sections show that the turbulent flow increases the average settling velocity of the inertial particles. This increase is due to preferential sweeping



**Fig. 16.** Effect of the blocked lateral motion on the settling rate. The simulations have been performed excluding particle motion in the horizontal direction (perpendicular to gravity).

(Wang and Maxey, 1993) which dominates other mechanisms that may slow down the settling rate, such as vortex trapping or loitering. An examination of the relative importance of the different mechanisms in a three-dimensional unsteady flow with a large number of particles is challenging. Recently, Good et al. (2014) performed numerical simulations in which the two mechanisms (i.e. preferential sweeping and loitering) were separated. This was done through blocking the particle motion in the horizontal direction, i.e. in direction perpendicular to the gravity. Blocking of the horizontal motion prevents natural tendency of the particles to sweep in the downward flow regions and therefore restricts the particleeddy interaction only in the vertical direction. They observed a reduced settling velocity of inertial particles in such conditions. In the present study we follow this idea. Four series of DNS simulations have been performed at four different values of turbulence intensity. Fig. 16 shows the difference between the average settling velocity and the associated terminal velocity (computed using the Stokes drag) for a given particle size. The different lines correspond to different values of energy dissipation rate. Several conclusions can be drawn from Fig. 16. First, the average settling velocity is lower than the terminal velocity for all particle/flow settings. This reduction is expected as the mechanism of preferential sweeping is disabled. Second for the droplets with radii smaller than 27.5µm the reduction is insensitive to energy dissipation rate. Third, the maximal reduction of the settling velocity takes place for droplets with radii  $\approx$  40  $\mu$ m. Presence of this minimum indicates that for large particles the sampling of U along the trajectory may have a dominant effect, as explained next. Rewriting particle equation of motion as  $\langle V \rangle - \tau_p g = \langle U \rangle - \tau_p \langle dV/dt \rangle$ and noting that the acceleration term on the right hand side is zero, it is evident that the reduction in particle sedimentation velocity depends strongly on selective behavior of  $\langle U \rangle$  at particle location. Namely, small particles ( $a < 27.5 \ \mu m$ ) experience a similar  $\langle U \rangle$  dynamics irrespective of the energy dissipation rate while larger particles sample a different fluid velocity on their trajectory causing the particle sedimentation response to change with  $\varepsilon$ .

# 5. Conclusions

Using direct numerical simulations, we have examined the different mechanisms and factors affecting the average settling velocity of inertial particles in a turbulent flow. The main focus is on water droplets of 10–60  $\mu$ m in radius suspended in air. It has been shown that the difference between the particle settling velocity and their terminal velocity is a roughly linear function of  $\sqrt{\varepsilon}$ . Simulations with the Stokes drag and various density ratios show that the average settling velocity is consistently larger or equal to the terminal velocity. This conclusion is in agreement with the recent study of Good et al. (2014). Additionally, we showed that the maximum increase in the settling velocity highly depends on the turbulence intensity or the energy dissipation rate at fixed  $R_{\lambda}$  and kinematic viscosity.

Furthermore, effects of the Taylor microscale Reynolds number on the settling rate have been investigated. Consistent with the observations of Yang and Lei (1998), we conclude that at a fixed energy dissipation rate and a low Taylor microscale Reynolds number ( $R_{\lambda} < 100$ ) the maximum level of increase scales with  $R_{\lambda}$  and hence u'. On the other hand, for larger  $R_{\lambda}$  the increase in settling velocity saturates and this saturation is observed only for lowinertia particles ( $a < 30 \,\mu$ m).

The DNS results were used to develop an empirical parameterization that relates the settling velocity to three nondimensional parameters, namely, *St*, *Fr* and  $R_{\lambda}$ . The accuracy of the new parameterization has been verified by comparing against DNS data. For a wide range of the nondimensional parameters we obtained quantitative agreement between the DNS results and the parameterization. Only for large  $R_{\lambda}$  and 1 < Fr < 20 the parameterization yields a rather qualitative agreement.

In order to examine the possible effects of the forcing mechanism, our simulations were carried out using two different forcing schemes. Results obtained using both stochastic and deterministic forcing method show a similar trend, although there is a significant difference in magnitude of the settling velocity for a given particle radius and  $R_{\lambda}$ . At a given grid resolution, particles settle faster in the flow forced by the deterministic scheme. This is partially caused by the different values of  $R_{\lambda}$ . Flow statistics show that the deterministic scheme allows a higher  $R_{\lambda}$  to be obtained. We expect the deterministic scheme to yield more coherent turbulent eddies. On the other hand the random nature of the stochastic forcing may affect the preferential concentration and consequently the settling velocity.

Additionally, we have examined the role of the gravitational acceleration. It has been shown that a larger gravity magnitude pushes the settling velocity of high-inertia particles (St > 0.2) to quickly converge to their terminal velocity. For low-inertia particles (St < 0.0504), however, we observe a small increase in the settling velocity when gravity is increased from 1g to 2g.

Another important issue explored in our study is the non-linear drag and its effect on particle settling rate. In the considered range of Stokes number and  $S_V$ , the difference in the settling velocity due to different drag forces is very small and is independent of the energy dissipation rate. Our results are in qualitative agreement with the results of Wang and Maxey (1993). We confirm that the non-linear drag slightly reduces the net increase in the settling velocity.

Finally, we performed simulations in which the horizontal motion of particles is blocked. Preventing horizontal motion turns off the preferential sweeping mechanism, which has a significant impact on the settling velocity of particles. Turning the preferential sweeping off causes the average settling velocity to be lower than the terminal velocity. Consistent with the recent study of Good et al. (2014) and Parishani et al. (2015) this implies that the dynamics of particle-eddy interaction in the horizontal plane plays a central role in influencing the particle velocity in the vertical direction.

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